



H2020 Grant Agreement No. 730562 – RadioNet

| | |
|-------------------------|---|
| <u>PROJECT TITLE:</u> | Advanced Radio Astronomy in Europe |
| <u>STARTING DATE</u> | 01/01/2017 |
| <u>DURATION:</u> | 48 months |
| <u>CALL IDENTIFIER:</u> | H2020-INFRAIA-2016-1 |
| <u>TOPIC:</u> | INFRAIA-01-2016-2017 Integrating Activities for Advanced Communities |



Deliverable 3.7

Update No 2 of the ‘Guide to the European ARC network’

| | |
|--------------------------|--|
| Due date of deliverable: | 2020-12-31 |
| Actual submission date: | 2020-09-16 |
| Leading Partner: | EUROPEAN SOUTHERN OBSERVATORY – ESO EUROPEAN ORGANISATION FOR ASTRONOMICAL RESEARCH IN THE SOUTHERN HEMISPHERE (ESO) |

Document information

| | |
|----------------------|--|
| Document name: | Update No 2 of the 'Guide to the European ARC Network' |
| Type | Report |
| WP | WP3 – Training |
| Version date: | 2020-09-16 |
| Authors (Institutes) | Evanthia Hatziminaoglou (ESO), Martin Zwaan (ESO), Gergő Popping (ESO) |

Dissemination Level

| Dissemination Level | | |
|---------------------|---|---|
| PU | Public | X |
| PP | Restricted to other programme participants (including the Commission Services) | |
| RE | Restricted to a group specified by the consortium (including the Commission Services) | |
| CO | Confidential, only for members of the consortium (including the Commission Services) | |

Index

| | |
|--|----|
| Guide to the European ARC Network' | 3 |
| 1 Purpose of this document..... | 3 |
| 2 ALMA in a nutshell..... | 3 |
| 2.1 The ALMA Regional Centres (ARCs) | 3 |
| 2.2 Organisation of the European ARC structure | 3 |
| 2.3 Role of the central ARC, the ARC nodes and the CoE | 4 |
| 3 How the ARC can help with your ALMA observing project..... | 5 |
| 3.1 The ALMA Science Portal..... | 5 |
| 3.2 The proposal stages | 6 |
| 3.3 The preparation of the Phase 2 material..... | 6 |
| 3.4 Tracking your project..... | 7 |
| 3.5 After data have been taken..... | 7 |
| 3.6 The data reduction pipeline..... | 7 |
| 3.7 Data reduction support..... | 7 |
| 3.8 Calibrated measurement sets | 8 |
| 4 Additional help from the European ARC..... | 8 |
| 4.1 Archive research support..... | 8 |
| 4.2 Need training? | 8 |
| 4.3 General questions..... | 9 |
| 5 How to organise a face-to-face visit..... | 9 |
| 5.1 Funding for visits..... | 9 |
| 6 The ARC nodes and CoE | 10 |
| 6.1 The Italian node in Bologna..... | 10 |
| 6.2 The Dutch node in Leiden: Allegro | 10 |
| 6.3 The German node in Bonn/Cologne | 11 |
| 6.4 The IRAM node in Grenoble..... | 11 |
| 6.5 The Nordic node in Onsala | 11 |
| 6.6 The UK node in Manchester | 12 |
| 6.7 The Czech node in Ondřejov | 12 |
| 6.8 The Portuguese CoE in Lisbon: PACE | 13 |

Guide to the European ARC Network'

1 Purpose of this document

The purpose of this document is to help users to find their way around the European ALMA Regional Centre (ARC). The European ARC is the 'one stop shop' for all European ALMA users. This document explains how the ARC structure works and how it can help users at the various stages of their ALMA observing project.

Updates of the guide are done yearly, prior to the Call for Proposals and reflect new procedures affecting the users or new services or tools provided to the user community by the network or the project. Minor updates can also come from the nodes, if something changes in their structure, expertise or services they offer. This document describes the latest update of the Guide to the European ARC Network done in March 2020 and includes the new service of the delivery of Calibrated Measurement Sets upon request, offered in the European ALMA user community since October 2019, as well as the change from the stale data policy to the raw data delivery. No other updates are foreseen in 2020.

2 ALMA in a nutshell

The Atacama Large Millimetre/sub-millimetre Array (ALMA) is expected to be the leading observatory at millimetre and sub-millimetre wavelengths for many decades to come. ALMA is a global collaboration involving East Asia, Europe, North America and the host country Chile. ALMA comprises 66 high-precision antennas equipped with receiver and digital electronics systems to observe in the frequency range from 30 GHz to 1 THz and at angular resolutions down to 6 milliarcsec. Using a fully dynamic scheduling system and innovative calibration strategies, the ALMA system allows observers to make the best use of the atmospheric conditions on the Chajnantor plateau at 5000 m altitude in the Atacama desert. The European contribution to the ALMA project is led by ESO on behalf of its member states.

ALMA has performed scientific observations since September 2011, in addition to regular commissioning and science verification. Each observing period offers another opportunity for astronomers to propose science projects with ALMA. Details on the new capabilities in this cycle are given in the Proposer's Guide, available at the ALMA Science Portal at <http://almascience.org>. More information on the ALMA project can also be found on the Science Portal.

2.1 The ALMA Regional Centres (ARCs)

The interface between ALMA and the astronomical community is provided by the three partners through the ALMA Regional Centres (ARCs). These are located at NAOJ in Mitaka, Japan, for the East Asian partnership, at ESO in Garching, Germany, for the European partnership, and at NRAO in Charlottesville, USA, for the North American partnership. The ARCs are staffed by scientists with expertise in radio astronomy, millimetre and sub-millimetre astronomy, and interferometry, and their purpose is to work with the community of astronomers to maximise the scientific productivity of the telescope.

2.2 Organisation of the European ARC structure

The ARC structure in Europe supports the scientific use of ALMA by the European user community. The ARC forms the interface between the ALMA observatory and the user community throughout the lifetime of a project, i.e., from proposal preparation to data analysis. In Europe, the services to the community are provided by a distributed network of ARC nodes at Bologna, Bonn/Cologne, Grenoble, Leiden, Manchester, Ondřejov, and Onsala, as well as an ALMA Centre of Expertise (CoE) in Lisbon, under the coordination of the ESO ARC in Garching.

The primary aim of the European ARC is to maximise ALMA's scientific return for European users. Each ARC node is staffed with scientists providing a range of experience in interferometry, (sub-)millimetre observing and ALMA data reduction and interpretation, as well as a range in scientific interests. The staff at the ESO ARC, the ARC nodes and the CoE work together to provide optimal support to users during the complete lifetime of a project from proposal preparation, preparation of the Phase 2 material, delivery of the calibrated science products to the users, and, if required, additional data reduction support, advice regarding observing strategies and can help with ALMA archival research. The ARC network organises community days, tutorials and workshops throughout Europe to reach out to the community and to train users in proposal preparation and data reduction. Several ARC nodes also supervise Master's and PhD students and give university lectures on ALMA-related topics.

The central point of contact for current and prospective ALMA users is the ALMA Helpdesk, which is available to all registered users from the *ALMA Science Portal*. There is one Helpdesk for all ALMA users, but queries from European users will be primarily addressed by the European ARC, either by scientists in Garching or by staff in one of the ARC nodes. The Helpdesk includes a library of “Knowledge-base” articles that address a number of common issues and questions. If the answer is not found in the suggested Knowledgebase articles, the user can submit a ticket. During normal operations, users can expect a response within two business days. During the week prior to a proposal deadline, extra staff will be assigned to the Helpdesk and every effort will be made to address time-critical questions in a rapid manner.

For all ALMA questions: use the ALMA Helpdesk to contact the ARC. The helpdesk is available through the [ALMA Science Portal](#)

2.3 Role of the central ARC, the ARC nodes and the CoE

The central ARC in Garching and the ARC nodes distributed throughout Europe work together to provide optimal support to European ALMA users. It is however important to realise that the roles that the central ARC, the nodes and the CoE fulfil are very different. For example, all face-to-face support, one of the core functions of the ALMA project, takes place at the ARC nodes.

The different roles of the ARC and the ARC nodes are outlined below. Note that the typical ALMA user does not need to know the specifics of how tasks are divided between the central ARC and the nodes: All initial contacts between the user and the ARC staff happen through the ALMA Helpdesk. The ARC staff may answer the user's question immediately, refer to documentation, or forward the query to an expert within the ARC-nodes network. The ARC may advise that face-to-face help is required, or the users may indicate themselves that they need to get in touch with their Contact Scientist (CS) or request face-to-face support.

The main tasks of the central ARC, of direct relevance to ALMA users, are:

- Phase 1 operations: distribution of the call for proposals to prospective European ALMA users, user support during proposal preparation as well as assistance in coordinating the proposal reviewing process.
- Phase 2 operations: manual generation of Scheduling Blocks when necessary, inspection of the automatically generated Scheduling Blocks and implementation of changes following successful change requests.
- Data product support: quality assurance (QA2) of science data sets and delivery of the final raw and reduced data to the PIs.
- Delivery of calibrated measurement sets.
- Staging of proprietary raw data deliveries.
- Archive operations: the ESO ARC node holds a complete synchronised copy of the ALMA archive.
- Running the ALMA Helpdesk.

The relevant tasks of the ARC nodes are:

- Quality Assurance (QA2) of science data sets.
- User support with proposal preparation by means of face-to-face meetings, dedicated workshops and through the ALMA Helpdesk.
- Phase 2 operations: assisting users with the validation of their Phase 2 material and answering questions on technical details pertaining the scheduling and execution of the proposed observations.
- Face-to-face help with data reduction, including expert support in data processing for specialised observing techniques.
- Help in archival research, including assistance to users of the ALMA archive in identifying and using the data products suitable for their scientific projects.
- Facilitate science coordination between ALMA users, i.e. providing space for small meetings, putting people with similar interests in contact if mutually desired etc.
- Close interaction with the community and regular updates on ALMA (Community Days and other events, newsletter, webpages), as well as community education through lectures, tutorials, etc.
- Outreach to the general public.

The tasks of the ALMA Centres of Expertise:

European ALMA CoEs are an addition to the European ALMA support structure since 2014. A CoE has a temporary status and may eventually acquire the status of full ARC node, once the conditions described in detail in the document “European ALMA Centres of Expertise: Tasks and Responsibilities” have been satisfied. The tasks of the CoEs are similar to those of the ARC nodes, with the exception that they do not provide CSs or face-to-face support for data reduction. Once a CoE becomes an ARC node, its role and responsibilities will fully align with those of the other ARC nodes.

The following section describes in more detail the kind of support to be expected throughout the lifetime of an ALMA project.

3 How the ARC can help with your ALMA observing project

For each scheduling period, a Call for Proposals for ALMA is issued. The ESO ARC distributes this call among the European user community by email and makes it available on the ALMA Science Portal (<http://almascience.org>). The call provides information about the available capabilities for each cycle, necessary information for the electronic submission of proposals as well as links to documentation such as the Proposer’s Guide, the ALMA Technical Handbook and the Observing Tool (OT) User and Reference Manuals, as well as the OT Quick Start Guide. For the complete list of documentation, see the relevant section on the ALMA Science Portal.

3.1 The ALMA Science Portal

All ALMA users need to register to the ALMA Science Portal. This is a single sign-on gateway to the various tools described below, the Helpdesk, the ALMA Science Archive, and other relevant information. Each astronomer who uses ALMA is assigned a single ARC for user support at the time they register. ALMA users should have *one* ALMA account, that they should update in case any of their profile details changes (e.g. e-mail address, affiliation etc).

To submit a proposal, download ALMA software, or gain access to the Helpdesk, you will need to register with the [ALMA Science Portal \(almascience.eso.org\)](http://almascience.eso.org)

3.2 The proposal stages

The Phase 1 material of the proposals must be prepared with the ALMA OT, available from the ALMA Science Portal, section "Tools". The OT requires the users to provide a scientific and technical justification, target specification, and sensitivity. The OT provides the users with time-on-target plus overhead specification, integration time estimation, transparency and atmospheric conditions, etc.

Need help with proposal submission? Use the [Helpdesk](#). Need face-to-face help for preparing a complicated proposal? Use the [Helpdesk](#) to coordinate a visit to an ARC node.

The ALMA Observing Support Tool (OST), hosted at the UK node (<http://almaost.jb.man.ac.uk/>) provides an easy web-based interface to preparing simulations. The user can upload a FITS image and 'observe' it with ALMA at different frequencies, bandwidths, configurations, integration times and weather conditions. The results are images as ALMA would produce them, together with point spread functions and maps of the uv -coverage. This information can help users in planning their ALMA observations.

More experienced users may want to use simulation capabilities included in the CASA data reduction package, which have more flexibility. For more details see <http://casaguides.nrao.edu>.

Staff at the EU ARC network assist users with the preparation of their Phase 1 material. This assistance covers help with the OT, the ALMA exposure time calculator and the simulator. Staff are also available for consultation regarding observing strategies.

Users may also choose to visit an ARC node for face-to-face support during proposal preparation.

After the submission deadline, all proposals submitted during the Main Call are ranked by the ALMA Proposal Review Committee (APRC), which makes a recommendation to the ALMA director on the scheduling of ALMA projects for the coming period. Proposals submitted during the Atacama Compact Array (ACA) stand-alone Supplemental Call are ranked through a distributed peer review process.

3.3 The preparation of the Phase 2 material

PIs are responsible for the submission of their Phase 2 Science Goals (SGs). PIs will be notified by email on the ranking of their project as well as on the timeline for the submission of their Phase 2 material. PIs can request assistance with the validation of their SGs from their CS, that will be assigned to all accepted A-, B-, and C-rated projects. Minor mistakes on the positions, velocities or frequencies can be corrected by the PIs directly. Major changes that might change the scientific content of the program or clash with another approved project, need special permission and should be requested via a special Helpdesk ticket. PIs are not requested to submit their validated SGs for the ACA stand-alone Supplemental Call.

Immediately after the Phase 2 deadline, the Scheduling Blocks (SBs) will be automatically generated by the observatory and will be released to ALMA operations in Chile for scheduling and eventual execution. No further alterations can be made without special approval, obtained via Helpdesk in consultation with a CS. Note that observers do not travel to Chile or elsewhere to take part in the observations - all ALMA observations are scheduled dynamically and executed in service mode.

The SBs can be executed at any time without further intervention from users, CS or ARC staff. If a problem is found at run-time, execution of the SB is halted. The Phase 2 team at the ESO ARC will fix the problem and the affected SBs are put back in the queue for scheduling.

Need the help of a Contact Scientist for preparing the Phase 2 material? Contact the ARC through the [Helpdesk](#).

Please note that in general PIs should submit their Phase 2 SGs as rapidly as possible to avoid any delays. In case of unavailability, PIs can delegate the generation of Phase 2 material to a person of their choice via the 'Project delegation' tab in their ALMA Science Portal User Profile. Projects for which the Phase 2 material is not successfully submitted by the deadline given in the time allocation email will be downgraded (note that this only applies for the main call, not for the ACA stand-alone Supplemental Call). The Phase 2 team will make sure all SBs will run smoothly at the telescope (e.g. check calibrator variability, target proper motion).

3.4 Tracking your project

The Snooping Project Interface (SnooPI) allows PIs to follow the status of their ALMA observing projects. It allows the PIs to search for their active projects and get an overview of the SBs that have been observed, those in the queue, those that are under processing, etc.

Follow project progress with [SnooPI](#), which is available on the Science Portal under the 'Observing' menu.

3.5 After data have been taken

ALMA observing projects are dynamically scheduled, taking into account weather conditions, the configuration of the array, the proposal ranking and possibly other constraints. Short projects may be observed within one day, whereas others may take several weeks or months to complete. After the observations for a project have been completed, users are notified by email. Note that B- and C-rated projects will not be carried over to the next cycle when not completed, while A-rated projects will automatically be carried over, for one cycle. Accepted proposals under Director's Discretionary Time (DDT) will stay in the queue for the 365 days following their approval.

All data pass through a multi-tier quality assurance (QA) scheme before they are released to the user. ALMA data sets can vary enormously in size depending on correlator settings, averaging times and total integration time. Data size may range from tens of GB for the shortest observations to many TB for lengthy projects. Users can download the data from the ALMA archive.

It is important to realise that *all* ALMA data are kept in the archive. All requests for data of any observing program are handled through this archive. This ALMA Science Archive is accessible through the Science Portal.

PIs can request via the ALMA Helpdesk that QA0-pass raw data should be staged for download as soon as they are available (i.e. before the full QA2 products are ready). The proprietary period for the entire MOUS will start after the raw data from the first Execution Block (EB) in the MOUS have been made available to the PI. The release of raw data to the PIs will have no implications on the QA2 process. The access of QA0-pass raw data replaces the previous stale data and early release of ToO and time critical data policies but with immediate start of the clock for the proprietary time.

3.6 The data reduction pipeline

When a block of data for a particular program has been observed, the ALMA data reduction pipeline is activated. This pipeline runs on dedicated machines in Chile and in the ARCs and it uses a heuristics system to automatically edit, calibrate, and image data. The pipeline heuristics have captured the knowledge of experienced radio interferometer users and incorporated it into the reduction scheme. The output of the data reduction pipeline is stored in the ALMA Science Archive and is subject to a final QA2 before the reduced data are made available to the user.

Pipeline calibration has been fully commissioned for most observing modes employed in the previous cycles, but about 20% of observing projects will still be calibrated manually using standard scripts (e.g., high frequency bands, full polarization). The data processing will be done using a combination of the pipeline and manual analysis using CASA.

3.7 Data reduction support

The data reduction pipeline produces high quality science products. Expert hands-on help with data reduction may be required however, especially when more complicated observing techniques are used. The ARC is the primary contact point for users who need assistance with data reduction. ARC staff provide basic data analysis support, ranging from simple advice, to providing appropriate data analysis documents and products, to detailed assistance for users who require it.

The primary responsibility of the ESO ARC is to ensure that the reduced data from all ALMA observing modes meet the requirements set by the approved observing proposal. The ARC nodes can provide further assistance to users who require it. A likely scenario may be the following: the ALMA data reduction pipeline

has processed a user's data set. If the data meet the requirements defined in the proposal, the data are delivered to the user. However, the user may wish to make images with higher spatial resolution or higher sensitivity. The ARC nodes can help to reprocess the data, using different weighting schemes in the imaging stage. Another user may need higher dynamic range image cubes and can work with an ARC node to apply self-calibration to the data. For the use of nonstandard analysis techniques and the utilisation of the advanced algorithms and software developed by a specific ARC node, ALMA users may also arrange a visit to this ARC node, where the best possible support can be provided.

The ARC nodes provide the latest (or any relevant) version of CASA, installed on high performance workstations with ample disk space. For any first line CASA support, please visit the CASA Guides website at <http://casaguides.nrao.edu>.

3.8 Calibrated measurement sets

Scientists have the opportunity to request calibrated data products. These requests are done through the ALMA Helpdesk: select Archive and Data Retrieval department followed by the "Data request" sub-category. Per ticket, scientist can request the calibrated data for up to 10 datasets (MOUSs), after which scientists in the ARC will stage the calibrated data products for download for a period of 28 days. The service is open both for ALMA PIs or Delegees with proprietary ALMA data and for archival users wanting to use datasets for which the proprietary time has expired. More information is provided at [this link](#).

4 Additional help from the European ARC

4.1 Archive research support

A complete, synchronised mirror of the ALMA archive is kept at ESO and is a valuable resource for data mining. Assistance for archive research encourages broader approaches to scientific investigations. Therefore, scientists in the ARCs, ARC nodes and CoE provide support for astronomers accessing the ALMA science archive. The ALMA Science Archive is available through the User Portal.

Basic archive research support: use the [Helpdesk](#). Advanced archive exploitation support: visit an ARC node (use the [Helpdesk](#) to coordinate a visit).

Prospective ALMA users are likely to want to consult the Archive when planning observations as well as for data mining. To ensure that the archive (and ALMA) is exploited to its full potential, ARC nodes or CoE staff can assist users in using the archive. This includes face-to-face visits. Note that only PIs and Data Delegees can access projects completed within the last 12 months.

The ARC nodes and CoE can also help users extract information effectively (such as potential calibration source properties, or whether there are public domain data matching the required criteria) as well as assist with analysis of archive data.

4.2 Need training?

The European ARC regularly organises workshops, tutorials and schools related to ALMA. Dedicated CASA data reduction tutorials are held throughout Europe, as well as sessions describing the ALMA OT and more general workshops explaining the current ALMA capabilities. The EU ALMA newsletter and announcements provide a list of all scheduled activities and the mailing lists of the ESO ARC and the local nodes are used to disseminate this information.

4.3 General questions

For any general questions related to ALMA capabilities, availability of observing modes, or other relevant topics, please use the Helpdesk.

Any questions related to ALMA: use the [Helpdesk](#).

5 How to organise a face-to-face visit

The preferred way to obtaining any type of support is through the ALMA Helpdesk: select “Submit a Ticket”, then “Face to Face Support (EU)”. In the Helpdesk ticket and in order to make the most of the face-to-face visit the user should give as much information as possible on the purpose of the visit and the proposed node and dates.

Support staff at this node are responsible for arranging the details of the visit. Each visitor is assigned a dedicated member of staff for support purposes and can expect that this support person will be able to respond as quickly as possible to their requests for help, within reason.

Users normally visit their local node. If there is no ARC node associated with the country where the user is working, or if the user and the ARC (node) staff decide that specialist support is best provided by another node, a visit to another node can be organised.

ALMA data reduction will normally be performed using dedicated computing facilities such as those provided by the ARC nodes. The data volumes are too large for laptops in most cases, except for the later stages of image or spectral analysis, possibly using subsets of the data. Please consult ARC staff if you want to use your laptop.

To ensure data are downloaded locally and ready for the PI, the PI must give explicit permission for processed and raw data to be downloaded from the ALMA Science Archive by the ARC node staff via the “Project delegation” tab in their ALMA Science Portal User Profile. ARC nodes provide facilities for the user to copy data onto a hard disk drive. Data backup should be discussed prior to the visit, so that the user can bring appropriate storage devices. Some nodes can also support ftp of data provided the link at the user's home institute is fast enough.

ARC nodes keep a copy of the data the user has been working on for at least a few weeks after the end of the user's visit. This enables the node staff to address any post-visit questions which may arise.

After a visit, the user is requested to submit a feedback form on the service received, indicating, among other things, whether the goals of the visit were met.

5.1 Funding for visits

It is foreseen that most users will apply for funding from their local agencies for travel to one of the ARC nodes or CoE. Some ARC nodes have access to funds for visiting scientists. Please check the details in the next section or visit the ARC node web pages. Funding is also available through the RadioNet MARCUs (Mobility for ARC Users) networking activity for visits that will take place by 31 October 2020. Requests for MARCUs funding should be sent via [this form](#). MARCUs-funded visits are paid by the user's institute, which then is reimbursed by RadioNet. Note that filling the feedback form is mandatory for users supported by MARCUs funding.

If you receive funding from RadioNet MARCUs and/or useful assistance from a European ARC node or the ESO ARC, please include this in the Acknowledgement section of any paper based on ALMA data of the supported project.

6 The ARC nodes and CoE

Each node has its own area(s) of expertise. Users are encouraged to visit the individual ARC node web pages to obtain the most up-to-date information. Most ARC nodes also maintain email lists that are used to inform their local communities about activities and region-specific news items.

6.1 The Italian node in Bologna

The Italian ARC is hosted by the INAF-Istituto di Radioastronomia in Bologna. Support is offered with proposal preparation and submission, tracking of accepted ALMA projects, data reduction with CASA, and with ALMA archive mining, polarimetry and mm-VLBI with ALMA, our three areas of expertise.

Visit the Italian ARC node at <http://arc.ia2.inaf.it>

The Italian ARC also investigates new techniques of data reduction and handling and is heavily involved in the development of the ALMA science data archive. It organizes seminars, tutorials and workshops to inform and train the Italian community and to stimulate scientific discussion and collaborations. ARC personnel are involved in teaching and supervising students.

Multi-wavelength research, especially through the use of archives, is becoming more and more requested. The Italian ARC therefore also offers assistance with reduction and analysis of data taken with other interferometers.

The Italian ARC researchers are involved in several (sub-)millimetre scientific projects covering different topics (from galactic and local-Universe studies to high-redshift galaxies and cosmology) and offer general scientific support on their specific area of expertise for continuum and spectroscopic observations.

At the host institute, the ARC node can accommodate visitors for face-to-face support and visits. Visitors can connect to a dedicated computer server, to access the data and run the CASA data reduction software.

The ARC node is connected to the outside world through a high-speed optical fibre network, allowing fast data transfer (10 Gbit/sec). Outside users of ALMA and CASA can access the ARC node cluster facilities for a certain period of time.

No local funding is available for visitors, but they can apply for MARCUs support.

6.2 The Dutch node in Leiden: Allegro

Allegro, the ARC node in the Netherlands, is located at Leiden Observatory. The node employs approximately six postdocs and staff members who coordinate the node's activities. Allegro provides general user support for the ALMA community and offers expert help on high-frequency and long-baseline ALMA observations, mm-VLBI observations, and science analysis tools.

Visit the Dutch ARC node at www.alma-allegro.nl

The Allegro node supports users in all the steps, from ALMA proposal preparation through data delivery and data analysis. We offer assistance tailored to the needs of the user and organise activities for our user community. Allegro has guest desks available for visitors, who can make use of the dedicated high-performance computing facilities and connected fast-access storage system.

Allegro also contributes to the progress of the field within its own expertise areas. The node hosts expert knowledge in high-frequency and long-baseline observations with ALMA. By investigating techniques to overcome the impact of atmospheric effects, such as water vapour fluctuations, we drive ALMA towards reaching its full potential at the highest possible resolution and frequencies. The team has also partnered with the mm-VLBI group at Radboud University to develop Very Long Baseline Interferometry with ALMA, offered since Cycle 4 with the Global Millimetre VLBI Array (3mm) and the Event Horizon Telescope (1mm). In addition, Allegro develops science analysis tools, like ARTIST/LIME, to assist scientists in the analysis and interpretation of their observations.

6.3 The German node in Bonn/Cologne

The German ARC node is a collaboration between the astronomical institutes at the universities of Bonn and Cologne. The node's main facilities are located at the Argelander-Institut für Astronomie in Bonn, right next to the Max-Planck-Institut für Radioastronomie and a 30-minute commute from the University of Cologne. Four staff and five postdocs are associated with the ARC node, working on various support tasks. Areas of special expertise include the combination of interferometric data with single-dish data as well as advanced analysis and modelling of spectral line data. The node maintains and continuously extends the Cologne Database for Molecular Spectroscopy and offers a number of data analysis tools developed in-house. In close collaboration with the mm-VLBI group of the Max-Planck-Institut für Radioastronomie, the node also supports the use of ALMA in global mm-VLBI networks.

Visit the German ARC node at www.astro.uni-bonn.de/ARC

Visitors to the German ARC node can use the spacious ARC visitors' room equipped with workstations, switches for laptops and a conference phone, and a parent-child office, located next to the ARC visitors' room. The workstations offer fast connections to two high-performance servers dedicated to ALMA-related work. The servers run all the latest ALMA software necessary for the proposal stage, data reduction and analysis and for ALMA archive searches. Up to two visitors can be accommodated simultaneously.

In addition to user support, development and services to ALMA operations, the German ARC node reaches out to the local community with annual community meetings, a local newsletter and talks at seminars and conferences. The node's regular training opportunities include (remote) participation in two comprehensive hands-on lecture series on radio interferometry and radio astronomy in general, a CASA tutorial and ALMA-OT tutorials. In addition, some members of the ARC node participate actively in general public outreach and teacher education.

No local funding is available for visitors but eligible visitors can apply for MARCUs support.

6.4 The IRAM node in Grenoble

IRAM (Institut de Radio-Astronomie Millimétrique) is an international institute supported by the Max Planck Society (Germany), the CNRS (France), and the IGN (Spain). The institute is operating the 30-m Pico-Veleta antenna and the NOEMA mm-interferometer.

Visit the IRAM ARC node at www.iram.fr/IRAMFR/ARC

The IRAM ARC node activities are built on the in-house experience with user support and on the involvement of the institute in the ALMA design and construction phases (e.g. the development of the ALMA real-time Telescope Calibration software). The face-to-face support for ALMA uses the same procedures and infrastructures as those used for the NOEMA support. A computer room dedicated to data reduction is available and each project is assigned to a local contact. The ARC node involves 4 core members (3 staff and 1 postdoc), with additional help from other IRAM staff. The IRAM ARC node provides a good opportunity to use synergies between the IRAM instruments and ALMA, as, e.g. an access to the full sky for interferometric observations or to large-scale surveys with the 30-m antenna.

Funding is available for travel and accommodation, with the usual IRAM rules: one mission/one person per project. Funding is however limited to astronomers affiliated to one of the IRAM funding agencies.

6.5 The Nordic node in Onsala

The Nordic ARC node's main mission is to support ALMA users in Sweden and other Nordic and Baltic countries. The node is physically located at Onsala Space Observatory (OSO) near Gothenburg, Sweden. OSO is the Swedish National facility for Radio Astronomy and is hosted by Chalmers University of Technology. OSO has a long history in millimetre/sub-millimetre astronomy and presently runs the 20-m single dish telescope at Onsala and is a partner in the APEX sub-millimetre single dish telescope at the ALMA site in Chile.

Visit the Nordic ARC node at www.oso.nordic-alma.se

The ARC node has high capacity data links for downloading archival data and high-performance computer facilities to support ALMA data reduction. There is ample dormitory and guest room accommodation on-site to support data reduction visits and small workshops. The Nordic ARC node includes experienced millimetre wave observers and IT support.

The node is active in community outreach to promote ALMA in the Nordic countries. The node also organises meetings for the whole Nordic community to explain the science that can be done with ALMA and train users to plan, submit, and reduce ALMA projects. As well as giving general face-to-face support, the node has expertise in the areas of astrometry, robust phase self-calibration, multi-frequency synthesis, deconvolution algorithms and polarisation. The node also maintains several advanced data analysis packages. In collaboration with academic staff at Chalmers, there is also extensive scientific expertise on-site in the areas of radiative transfer and astrochemistry.

No local funding for travel and accommodation is available but eligible visitors can apply for MARCUs support.

6.6 The UK node in Manchester

The UK ARC Node is centrally located at the University of Manchester, Jodrell Bank Centre for Astrophysics, conveniently placed for all national and international transport links and high-speed internet connectivity. Five ARC staff (plus IT and administrative support) divide their time between face-to-face user support and related research and development. Facilities for visiting astronomers include a large, dedicated room with ample desk space and powerful computers, which can also be used for small meetings and tutorials. Absolute beginners or experts are welcome and resources can be made available by arrangement for teams working on ALMA projects with demanding computational and skills requirements.

Visit the UK ARC node at
<http://almadev.jb.man.ac.uk/>

The UK ARC Node draws on local and national expertise to offer specialised assistance in a number of areas:

- Interferometry skills, in particular advanced calibration, extended array configurations (high resolution), wide-band high-sensitivity continuum imaging (multi-frequency synthesis), mosaicking and wide-field imaging (including combining ALMA configurations to improve image fidelity), and polarimetry.
- Multi-instrument research, in particular comparing data from different telescopes, (e.g. *Herschel*, JCMT, e-MERLIN, EVN) combining ALMA and compatible data, and interoperability between CASA and AIPS. The UK ARC Node is part of the JBCA Interferometry Centre of Excellence which provides support for these instruments, others such as LOFAR and prospective SKA users.
- Data management, in particular data mining the ALMA Archive, development of pipelines and scripts for large projects, standards and documentation for data publication.
- Simulations using the OST (<http://almaost.jb.man.ac.uk/>) or inside CASA: choosing suitable input images, estimating the brightness on ALMA scales and so on.

Eligible visitors can apply for MARCUs support to visit the UK ARC node. Furthermore, PIs whose host institutions are unable to fund travel to the UK ARC Node from their consolidated grant, or other sources, can contact STFC (Science and Technologies Facilities Council) with an explanation. STFC can then look at awarding funding in exceptional circumstances and will pay for travel if they decide that it is justified.

6.7 The Czech node in Ondřejov

The Czech node is located at Ondřejov observatory near Prague, well accessible by public transport, hosted by the Astronomical Institute of the Czech Academy of Sciences. The node's mission is to provide support to ALMA users namely from the central and eastern regions of Europe, but due to its exclusive expertise in solar observations to all European users. The node's duties are distributed over two staff members and four postdocs, plus IT and administrative support.

Visit the Czech ARC node at
www.asu.cas.cz/alma

The team is ready to provide (face-to-face) support with preparation and submission of proposals to ALMA, with data processing, interpretation of observations and numerical modelling of some observed processes. A dedicated office equipped with computers and high-speed internet connectivity is available for up to three visitors. Affordable accommodation in the guest rooms of the Astronomical Institute at Ondřejov can be offered. The expertise areas of the Czech node are:

- Solar ALMA observing mode: specifics of proposal preparation, calibration and imaging; solar radio spectroscopy; non-LTE radiative transfer; physics of the solar chromosphere, prominences and filaments; solar flare plasma astrophysics.
- Galaxy evolution in clusters and groups, triggered star formation, proto-planetary disks, Galactic Centre and AGN.
- High-resolution microwave laboratory spectroscopy of small molecules, in collaboration with the University of Chemistry and Technology in Prague.

No local funding is available for visitors, but eligible visitors can apply for MARCUs support.

6.8 The Portuguese CoE in Lisbon: PACE

The Portuguese ALMA Centre of Expertise (PACE), operating since July 2014, is hosted by the Institute of Astrophysics and Space Sciences and located at the Astronomical Observatory of Lisbon. A team of radio astronomers devotes part of their time to PACE duties, such as organisation of events for community support, or face-to-face user support on proposal preparation and submission, as well as archival research.

Visit the Portuguese CoE at
<http://pace.oal.ul.pt>

Dedicated workstations, including access to a Nonius Cluster, are available to visitors.

The PACE astronomers have experience in mm-to-radio science. Their interests lie in the fields of galaxy formation and evolution, planetary atmospheres and archival research. The experience in these topics will be strengthened in order to improve the capabilities of the EU ARC network user support.

PACE also maintains a strong outreach team through which it educates the national community on mm-to-radio astronomy.

No local funding is available for visitors, but eligible visitors can apply for MARCUs support.

Copyright

© Copyright 2020 RadioNet

This document has been produced within the scope of the RadioNet Project. The utilization and release of this document is subject to the conditions of the contract within the Horizon2020 programme, contract no. 730562.

RadioNet has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 730562